

## 4.3 Newton's Universal Gravitation

November 16, 2017 12:31 PM

$$F_g = \frac{G M_1 M_2}{R^2}, \quad G = 6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2}$$

$F_g$  directly proportional to the masses and inversely proportional to the distance between the centers of the masses squared.

Ex pg 142

Practice pg 142 #1 a)  $7.5 \times 10^{-8} \text{ N}$  b)  $3.0 \times 10^{-7} \text{ N}$  }  $4F_g$

## Field Explanation

- Scientists invented the concept of fields to explain how a force can exist between 2 objects not in contact.

- a mass is surrounded by an invisible gravitational field and can interact with another mass' field.

- gravitational fields are vectors

gravitational field strength  $\rightarrow g = \frac{F_g}{m}$   $\leftarrow$  Force of gravity on a test object  
 $\leftarrow$  mass of the test object

$\rightarrow g = \frac{GM}{R^2}$   $\leftarrow$  mass producing the field  
 $\leftarrow$  distance from the centre of the field producing object.

Note:  $R$  is the distance from centre, so

"500km above surface of  $\oplus$ " means

$$R = (500\text{km} \times \frac{1000\text{m}}{1\text{km}}) + 6.38 \times 10^6\text{m}$$

$\nwarrow$  radius of  $\oplus$

## Weightlessness

- actually "apparent" weightlessness - zero force felt from supporting structures

ex: roller coaster  
descending in elevator  
running  
vomiting comet  
orbit (Constantly falling)

$$F_c = F_g$$

## Orbit

- speed needed to stay in circular orbit around  $\oplus$

$$F_c = F_g$$
$$\frac{m_s v^2}{R} = \frac{G M_\oplus m_s}{R^2}$$

$m_s$  = mass of satellite  
 $R$  = from satellite to centre of  $\oplus$

$$v = \sqrt{\frac{G M_\oplus}{R}}$$

- period of a satellite

circumference of a circle  $\rightarrow$

$$v = \frac{2\pi R}{T} = \sqrt{\frac{G M_\oplus}{R}}$$

time to go around once  $\rightarrow$

solve for  $T$

$$T = \frac{2\pi R^{3/2}}{\sqrt{G M_\oplus}}$$

$$\frac{T}{2\pi R} = \sqrt{\frac{R}{G M_\oplus}}$$
$$\frac{T}{2\pi R} = \frac{R^{1/2}}{\sqrt{G M_\oplus}}$$
$$T = \frac{2\pi R^{1/2} R^{1/2}}{\sqrt{G M_\oplus}}$$
$$= \frac{2\pi R^{3/2}}{\sqrt{G M_\oplus}}$$

once

solve  
for  
T

$$T = \frac{2\pi R^{\frac{3}{2}}}{\sqrt{GM_{\oplus}}}$$

$$= \frac{2\pi R^{\frac{3}{2}}}{\sqrt{GM_{\oplus}}}$$

- Escape Velocity
  - the speed needed to overcome the force of gravity.

$$E_k = E_p$$
$$\frac{1}{2}mv^2 = \frac{GMm}{R}$$

planet  
test object

$$v = \sqrt{\frac{2GM_{\text{planet}}}{R}}$$

## Gravitational Potential Energy

$$E_p = mgh$$
$$= m\left(\frac{GM_{\oplus}}{R^2}\right)R$$

$$E_p = \frac{GM_{\oplus}m}{R}$$

in text

$$E_p = -\frac{GMm}{R}$$

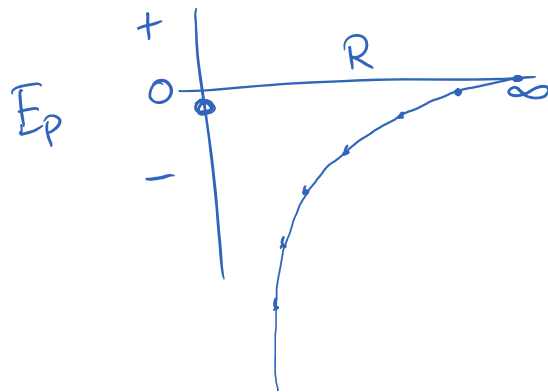
mass of planet  
mass of test object

Why is it negative?

At the surface  $E_p = 0$  (normal  $E_p = mgh$  calculations)

BUT - we define  $E_p = 0$  when 2 objects are infinitely far apart ( $E_p = \frac{GMm}{\infty} = 0$ )

- as objects get closer,  $E_p$  decreases
- $\therefore$  decreasing from zero, yields a negative #



Practice pg 143 # 2,3  
 pg 145 # 1-4  
 pg 146 # 1-2  
 pg 148 # 1-2